

European Southern Observatory

38<sup>th</sup> User Committee Meeting Garching, April 10-11, 2014

# **Observing Tools**

## Feedback from the COSMIC-LAB project

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INAF – Osservatorio Astronomico, Bologna (Italy) on behalf of the COSMIC-LAB team

Garching, April 11, 2014





## THE PROJECT



Star Clusters as Cosmic Laboratories for Astrophysics, Dynamics and Fundamental Physics

- ✤ 5-year project funded by the European Research Council (ERC)
- ✤ PI: Francesco R. Ferraro (Dip. of Physics & Astronomy Bologna)
- Co-Is: B. Lanzoni, A. Mucciarelli, E. Dalessandro + 4 Post-doc and 4 PhD
- kickoff: May 2011

#### AIM: understanding the complex interplay between dynamics and stellar evolution

+ HOW: using **globular clusters** as cosmic laboratories and

Blue Straggler Stars Millisecond Pulsars Intermediate-mass Black Holes as test particles





## **MID-TERM REPORT – first 30 months**

### **Granted telescope time:**



More than 200 orbits with HST and 500 hours at the 8-10m class telescopes have been assigned to projects related to COSMIC-LAB

✦ HST: Cycle 19 = 39 orbits Cycle 20 = 28 orbits Cycle 21 = 15+131 orbits



### **Published papers:**

24 papers have been published in peer-reviewed journals









#### web-page: http://www.cosmic-lab.eu/





www.cosmic-lab.eu



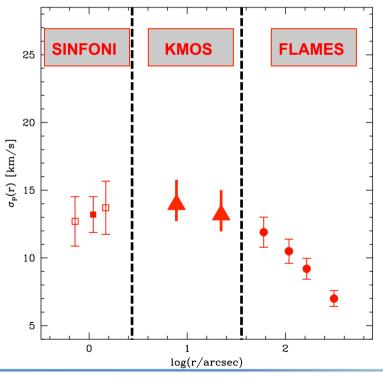
## Focus on VLT observations

### **Cosmic-Lab: Probing globular cluster internal dynamics**

Radial velocity dispersion and rotation profiles for 30 Galactic globular clusters ESO-VLT LARGE PROGRAMME (P93+94+95) KMOS + FLAMES PI: Ferraro 194 hours

Inner velocity dispersion and rotation profiles of five concentrated globular clusters from the radial velocities of individual stars

COMPANION PROPOSAL (P93) SINFONI PI: Lanzoni 31 hours

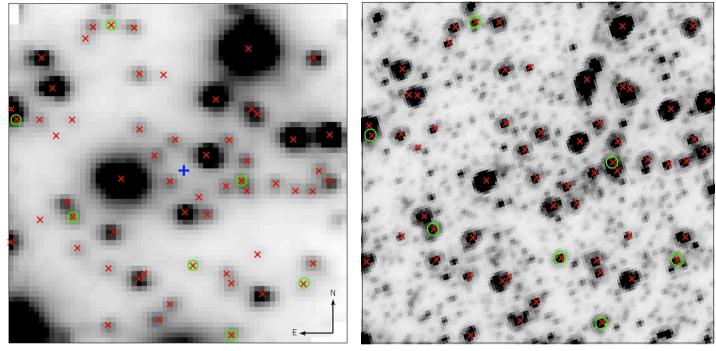




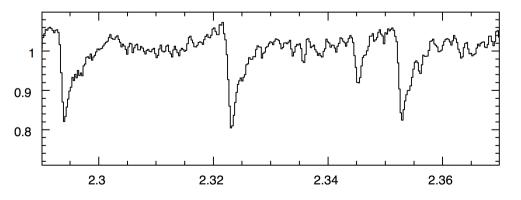
www.cosmic-lab.eu



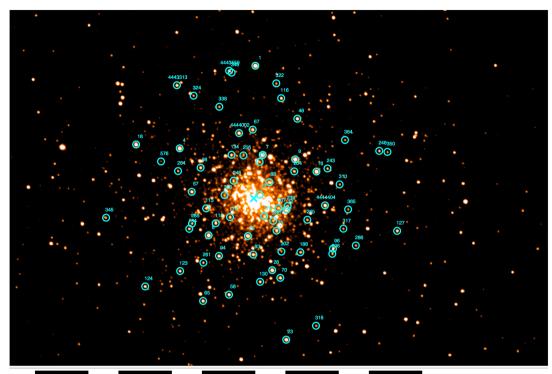
## VLT-SINFONI AO-assisted SINFONI HST/HRC



pilot project in the core
region of NGC 6388
K band spectra at R~4,000
~60 stars at r < 1.6 arcsec</pre>



## VLT-KMOS deployable IFU

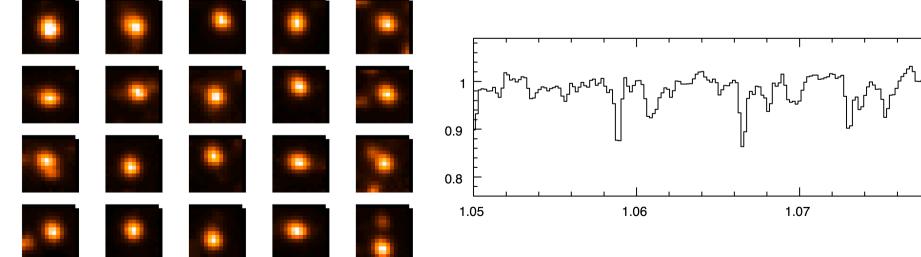


SV - pilot project in the NGC 6388 central region

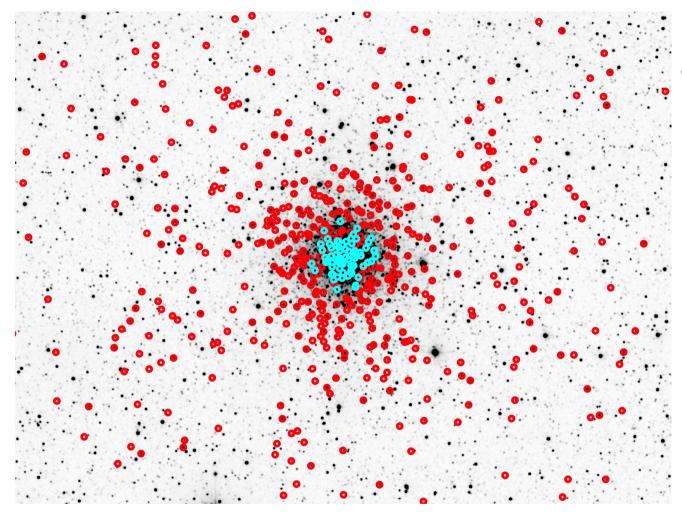
**YJ** band spectra at R~3,400

nod to sky mode

4 pointings, ~90 stars at R<70"







NGC 6388 outer regions

GIRAFFE/MEDUSA HR21 Cat I<17

## Focus on VLT observations

### **Cosmic-Lab: stellar systems with large pop of MSPs**

Bulge building blocks: chemistry and kinematics of Terzan 5, Liller 1, Terzan 6 ESO-VLT Proposals (P85-P91) X-Shooter + FLAMES PI: Ferraro 10 nights + 23 hrs

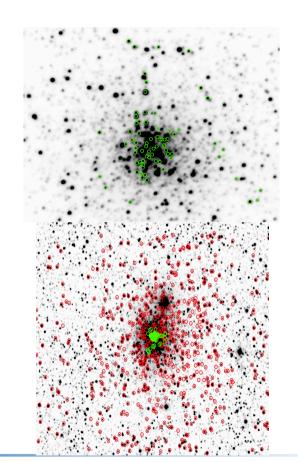
### high reddening $\rightarrow$ A<sub>v</sub>>7 mag

### XSHOOTER

VIS arm (RVs, CaT) slit 0.9", I<18 NIR arm (RVs, chemistry) slit 0.6", J<16

### FLAMES

GIRAFFE/MEDUSA HR21 (RVs, CaT), I<17







# Focus on VLT observations P2PP – general considerations

## *facts*/requests

almost every period new releases of P2PP and related instrument software/tools releases for Mac OS should become a standard

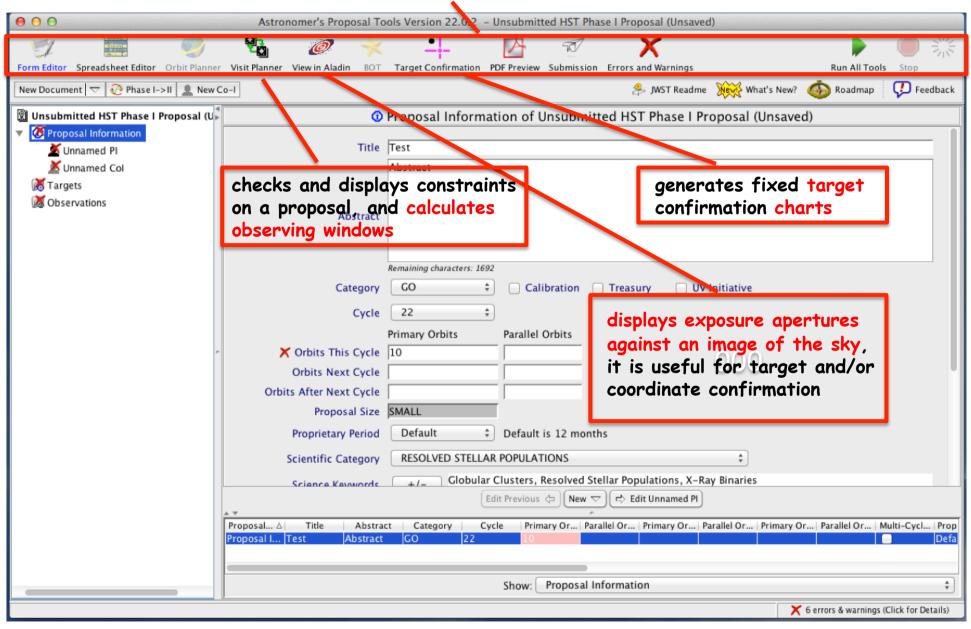
number of independent packages to manage is becoming large

**tools**, e.g. to produce finding charts, to check target visibility, to recover basic info (e.g. target and NGS coordinates, id, setups etc.) from Phase-I etc., **inside the P2PP** 

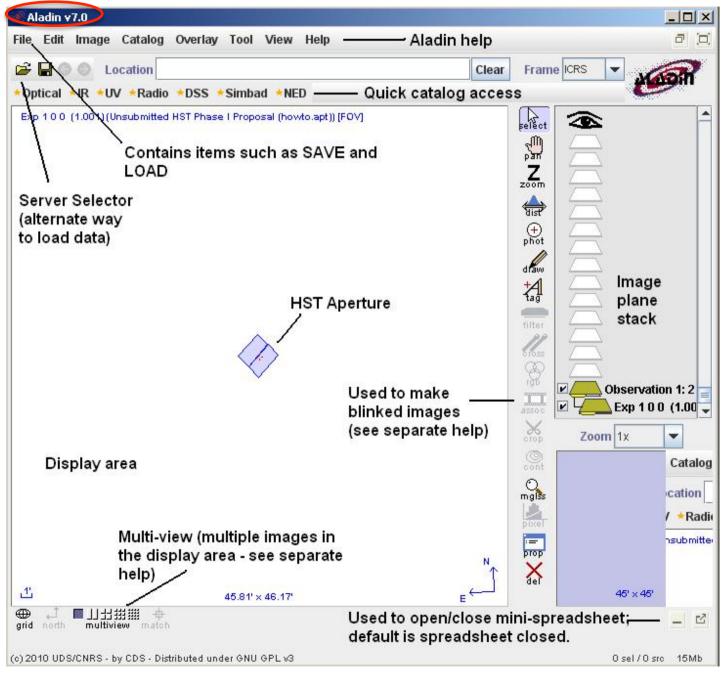
e.g., HST APT

## e.g., the HST APT

tool bar

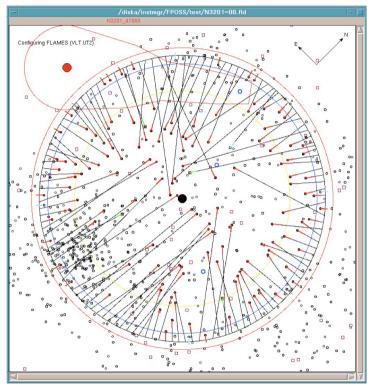


## e.g., the HST APT



## Focus on VLT observations: Positioners

### **FPOSS-FLAMES**



> successfully used by the Community since many years It does not run on Mac OS (only on a 32 or 64-bit Linux pc)

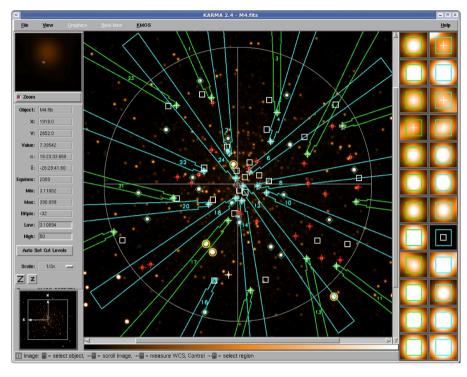
> The finding charts must be prepared with the FIMS software that runs only on 32-bit Linux pc

> It <u>does not allow</u> to load a FITS image and overplot the target catalog for an easier identification

> It does not allow to save intermediate configurations (e.g., with partial fiber allocation)

> Once a fiber configuration has been finalized (saved in a .ins file) it is not possible to update/modify it (e.g., change some fiber allocation)

## Focus on VLT observations: Positioners



### KARMA-KMOS

- It <u>runs</u> on both Mac OS and (32 or 64-bit) Linux pc
- > It <u>allows</u> to load a FITS image, and to overplot target positions
- > It <u>allows</u> to modify the display setup (e.g., colors, cuts etc.)

valuable if one could offset a given ifu from the nominal target position, e.g., to include an additional target that is just outside of the ifu fov, without having it automatically flagged as sky

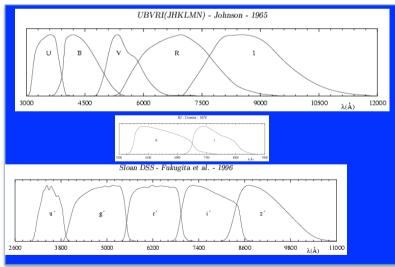
> to observe globular clusters in nod to sky mode, one must load a fits image with a large fov (>>KMOS) for checking offset positions (a few arcmin)  $\rightarrow$  ~ only 2MASS (low resolution, bad to check target allocation) valuable if one could have the option to load a high res image (small fov) to allocate/check targets, then change to 2MASS to allocate/check sky

## Focus on VLT observations: SINFONI

## ETC facts/requests

AO-mode requires R and (B-R) photometric info for the guide star can V and (V-I) be an alternative?

(MCAO WFS response curve not found) if R and (B-R) critical, in which photometric system? (e.g. Johnson, Cousin, SLOAN; Vegamag or AB etc.)



e.g., for a star with 
$$T_{eff}$$
= 5000 K  
 $\Delta(R_c - R_J) \cong +0.2 \text{ mag}$   
 $\Delta(R_c - R_S) \cong -0.2 \text{ mag}$   
 $\Delta(R_{Vega} - R_{AB}) \cong 0.4 \text{ mag}$ 

also, the observed (B-R) can be red since the star is intrinsically cool (BB-like) or reddened (~ $\lambda^{-1.5}$ ) or both

Focus on VLT observations: SINFONI ETC *facts*/requests

output: s/n and encircled energy over the DL core max intensity (obj+bg) and bg per pix & DIT

however, in crowded stellar fields

> the background can be enhanced due to the unresolved stellar continuum

the usable star signal is the one in the brightest pixel and occasionally around it (within about the PSF DL core)

#### it would be very valuable if the ETC could

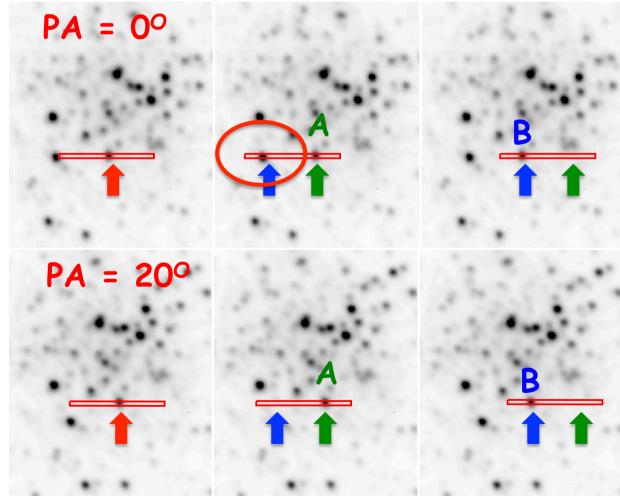
> allow to change the bg value

> provide also the s/n of the brightest pixel (subtracted by the bg), now it can be computed manually by the user

## Focus on VLT observations X-Shooter nod-on-slit tool inside the P2PP

to visualize target and the sky in the A and B positions within the slit, with varying the nodding throw and position angle

tool useful/available for any IR slit spectrograph



slit  $\rightarrow$  11" throw  $\rightarrow$  6" position angle  $\rightarrow$  0°

in A sky contaminated by a star

slit  $\rightarrow$  11" throw  $\rightarrow$  6" position angle  $\rightarrow$  20°

no contamination

# **Observing Tools**

## Feedback from the COSMIC-LAB project

## Summary

- > tools (e.g. finding charts, target visibility, etc.) within the P2PP
- new tools:

prop-info: to transfer Phase I info (targets coord, id, etc.)
into Phase 2 OBs (also some info into the Readme file)
nod-on-slit: to set optimal nodding throw and PA for IR
slit spectrographs (e.g. XShooter, CRIRES+)

- > some new features:
  - in FPOSS and KARMA
  - in the SINFONI ETC



